

Structure and Ion Composition of the Martian Wake Boundary Layer

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CESR, Toulouse, France

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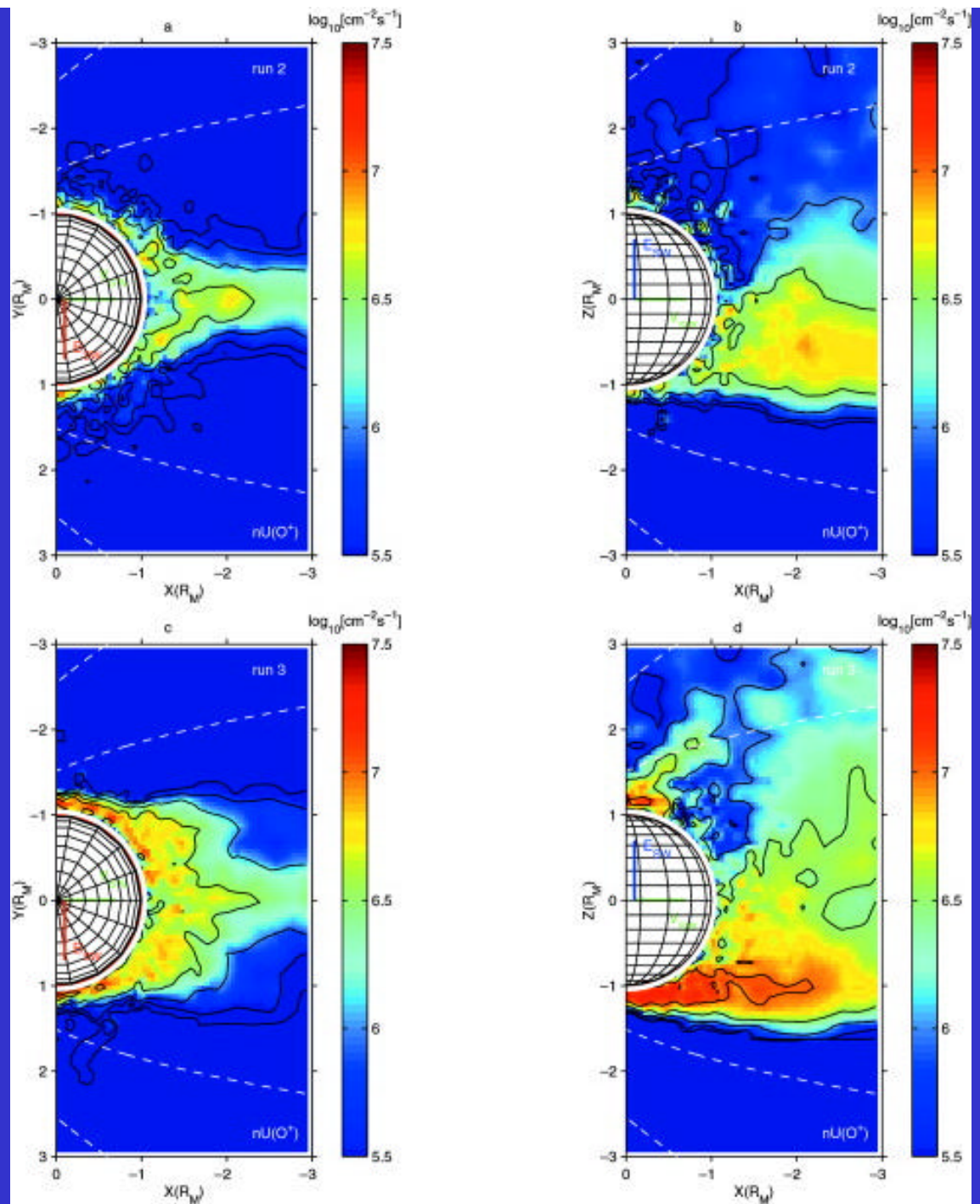
IRF, Kituna, Sweden

D. Winningham

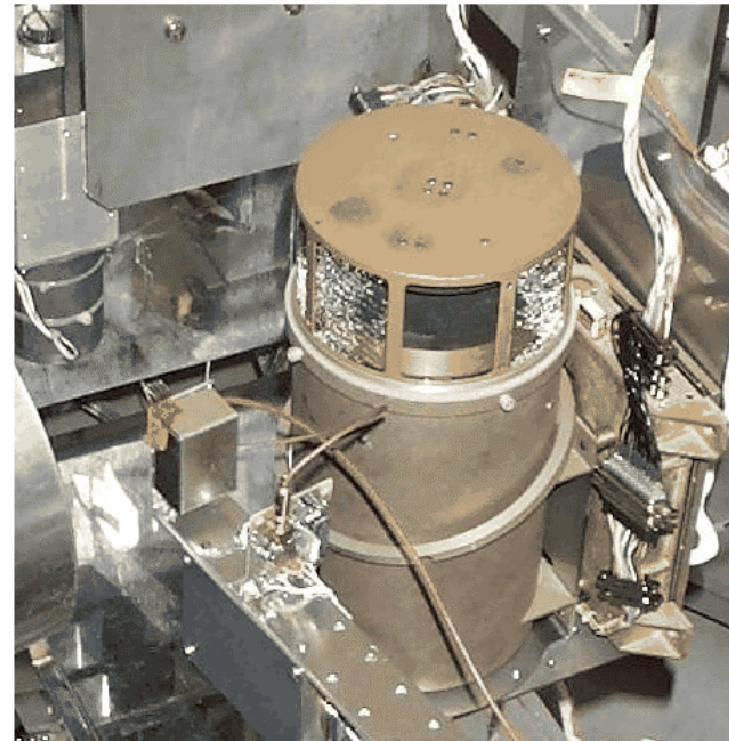
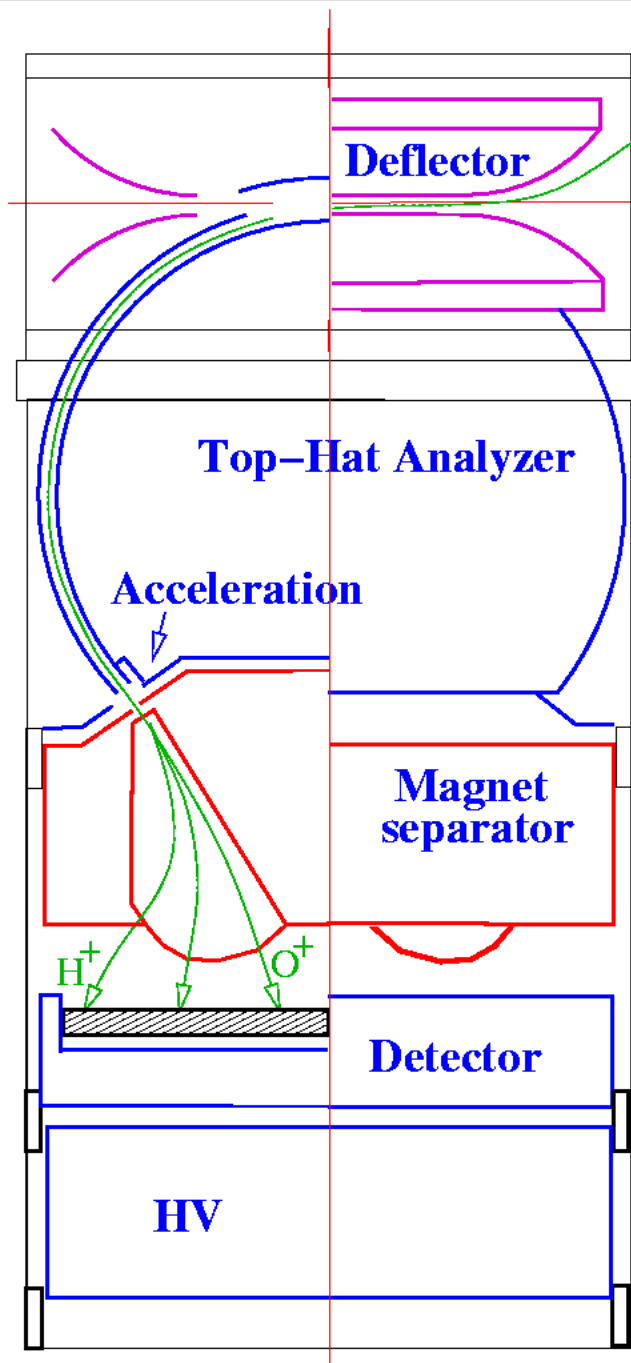
SWRI, San Antonio, USA

ASPERA Team

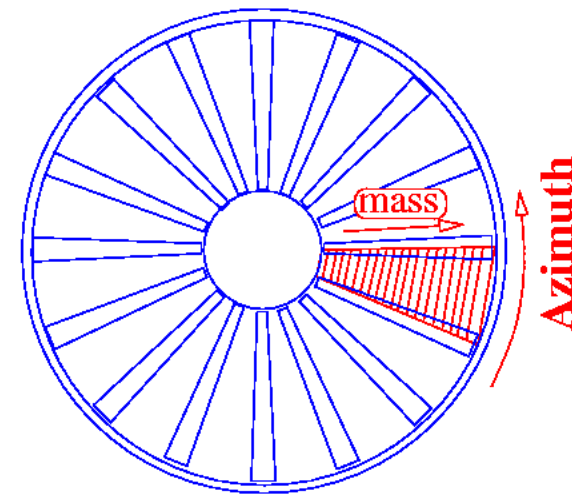
Ion Escape from Mars



Esa Kallio and Pekka Janhunen, JGR, V.107, p. 1035, 2002

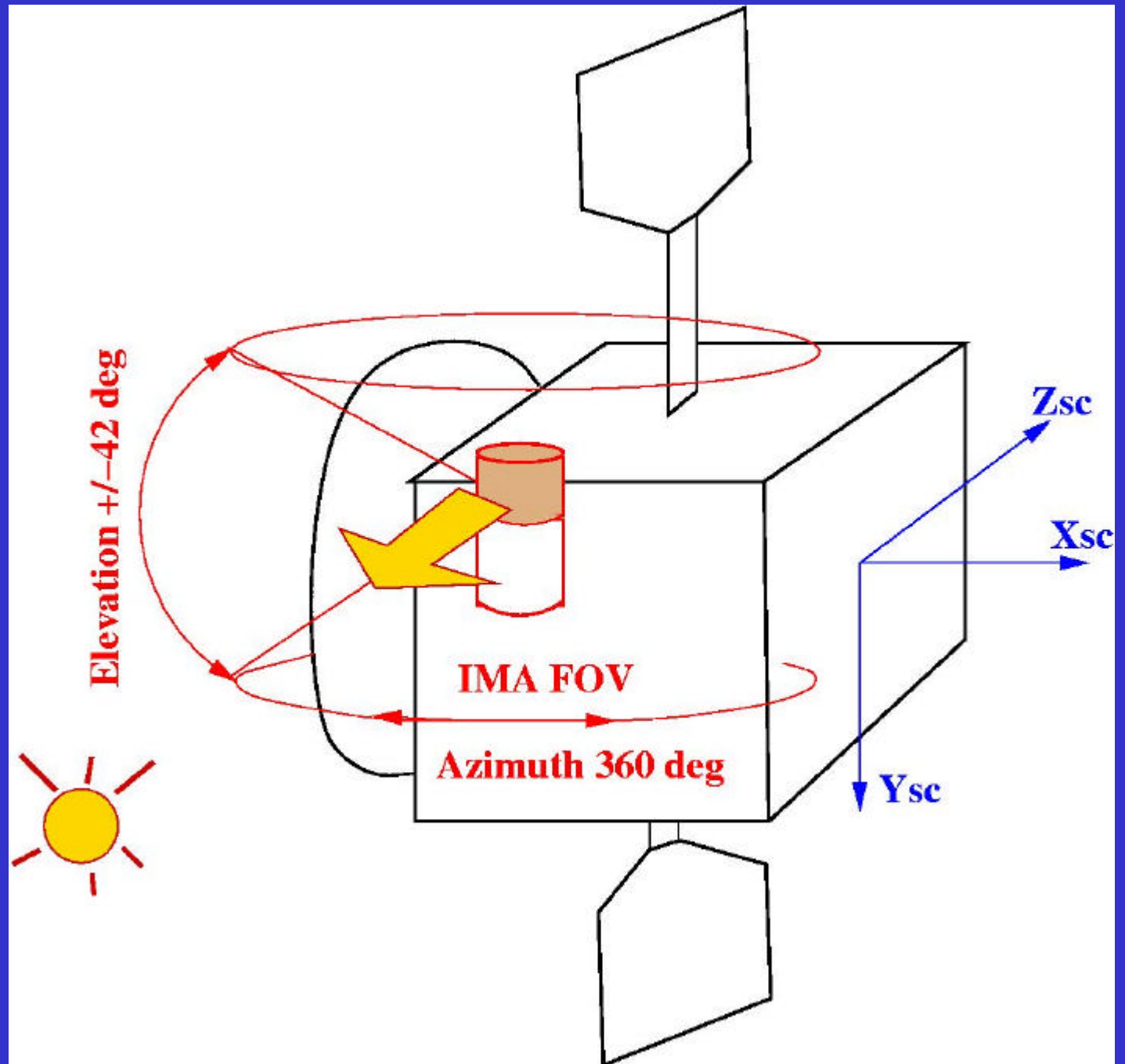


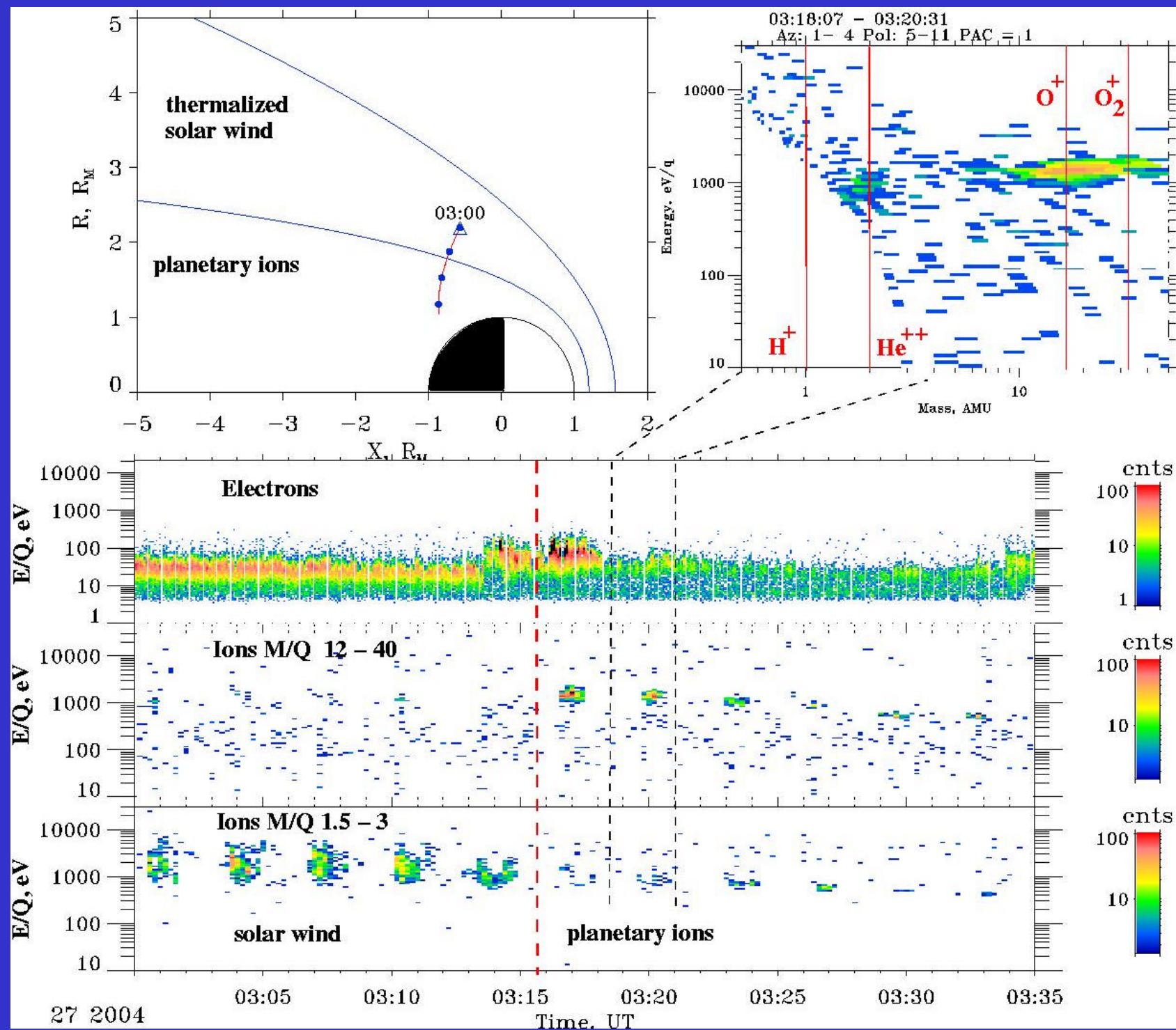
**16 azimuthal sectors
32 mass bins**

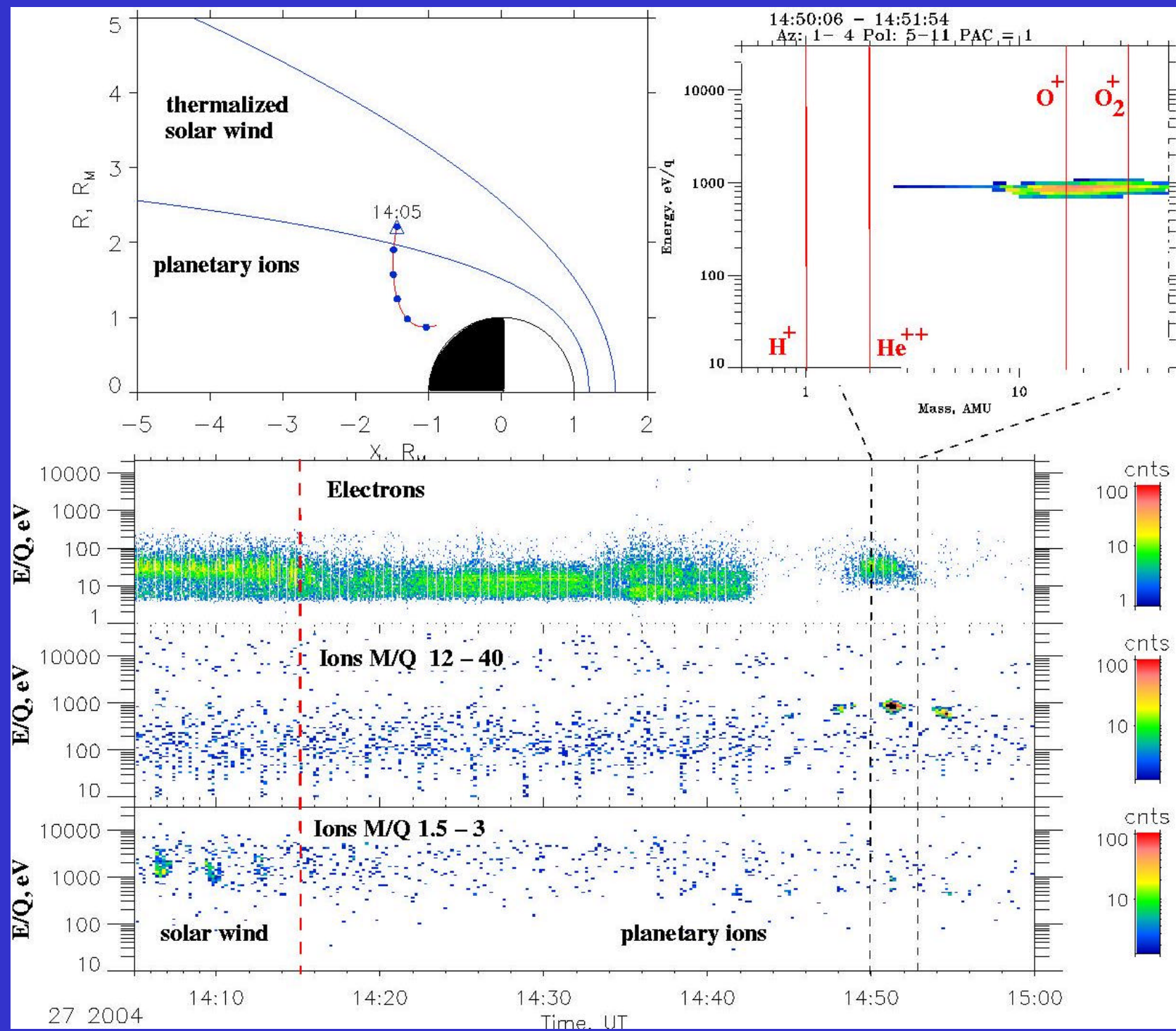


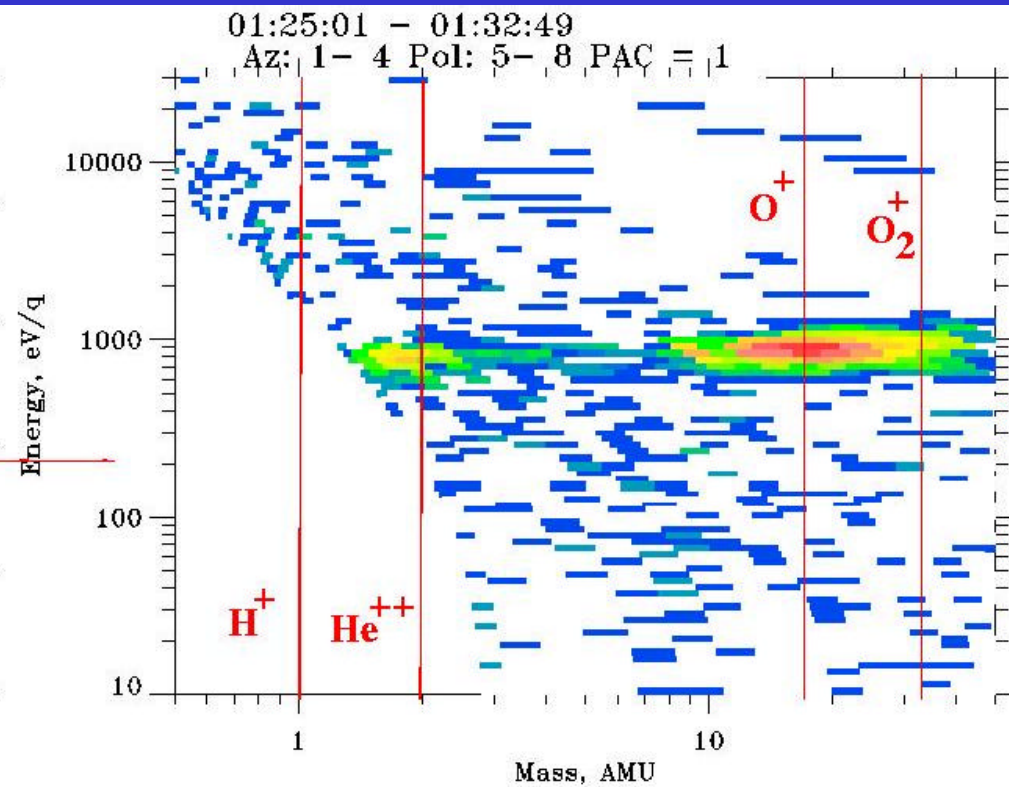
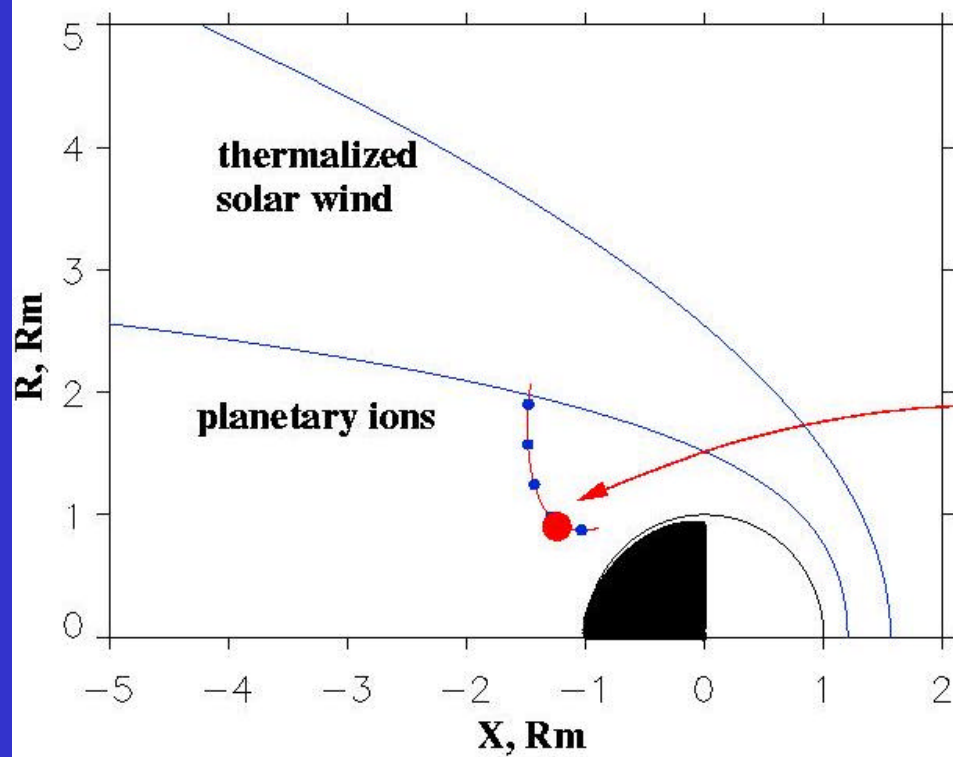
IMA mass-analyzer conceptual design

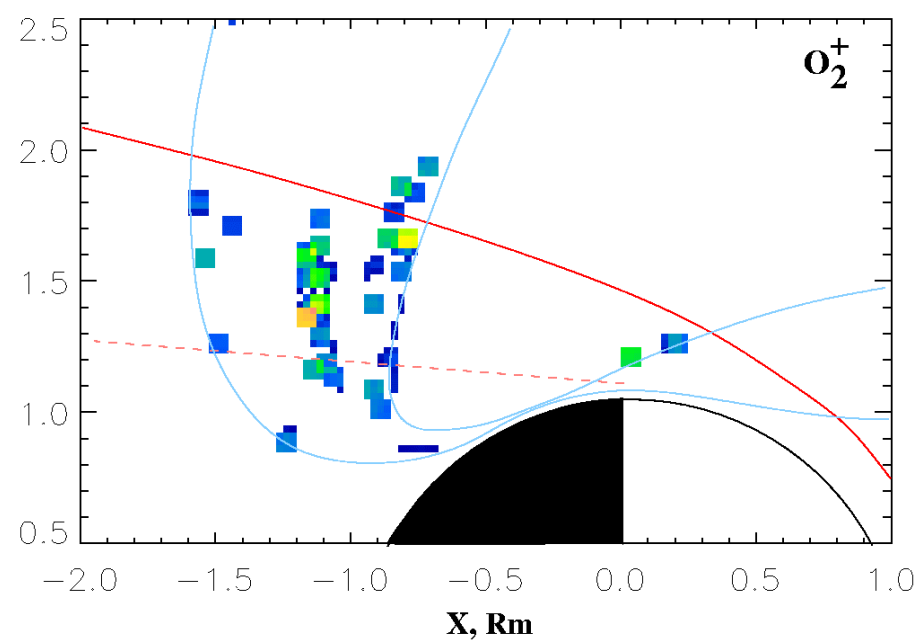
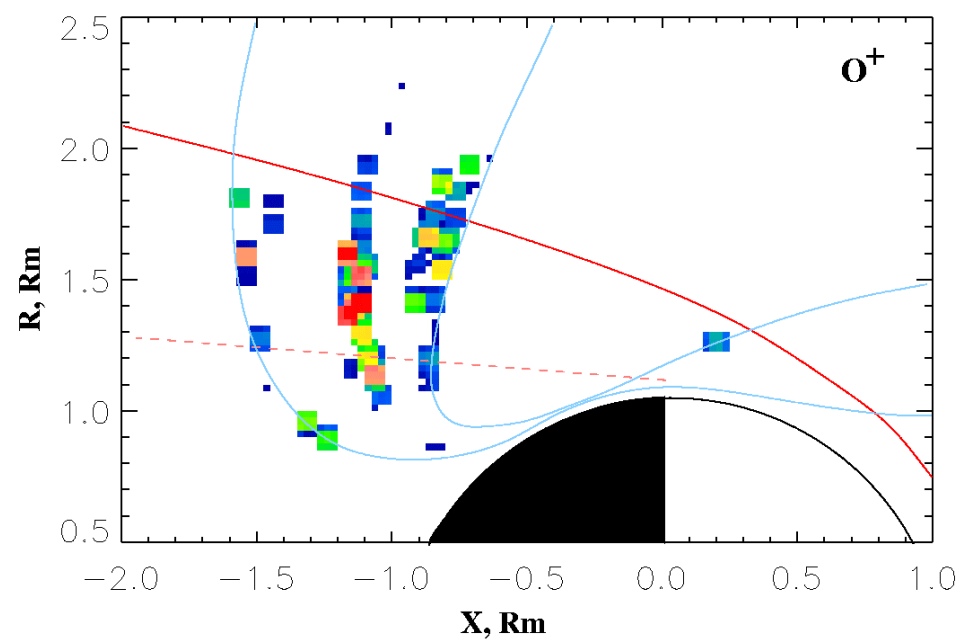
IMA attitude and Field of View

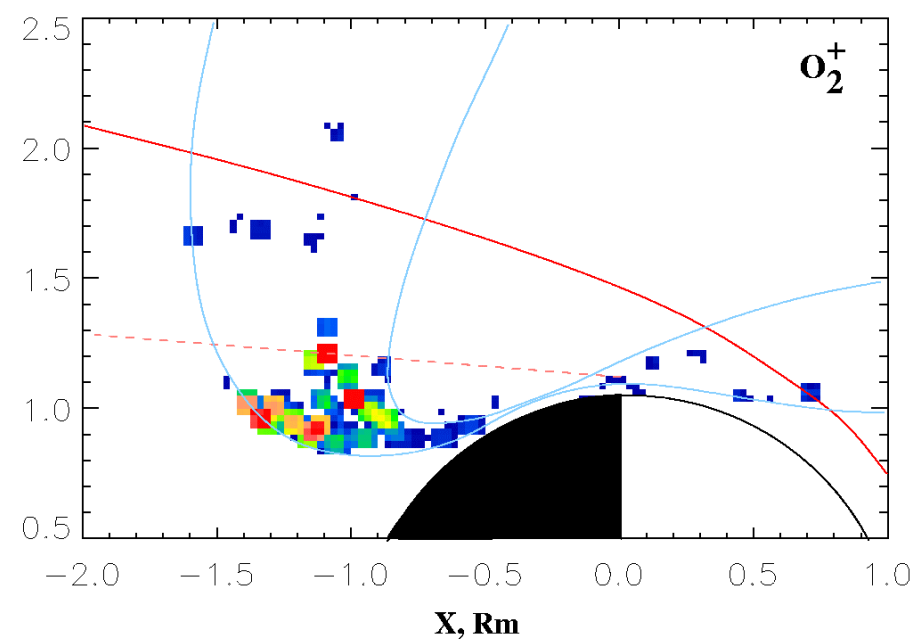
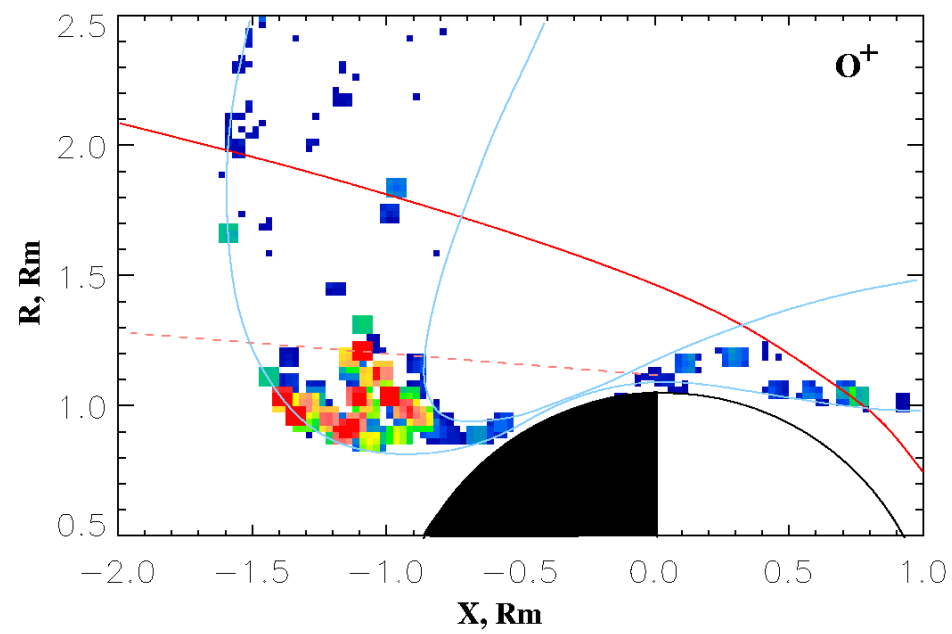


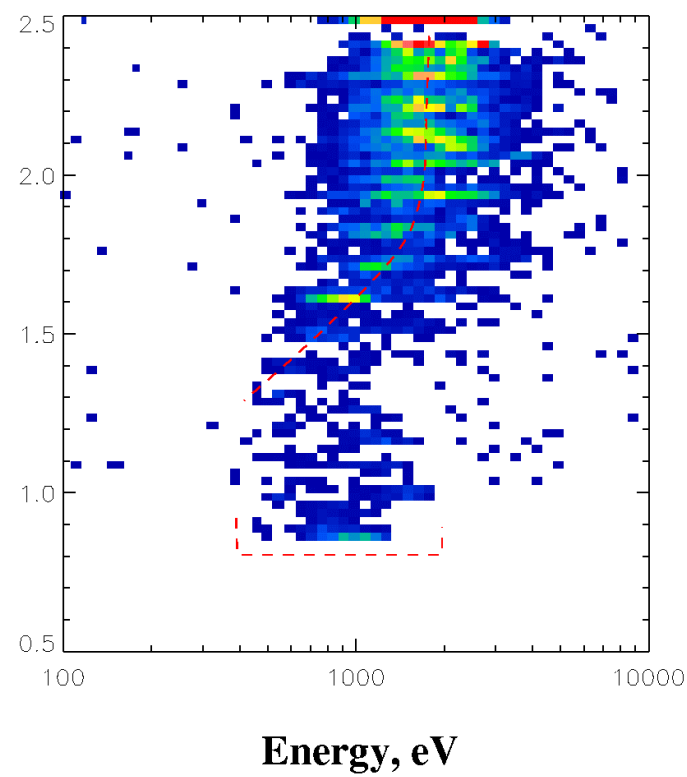
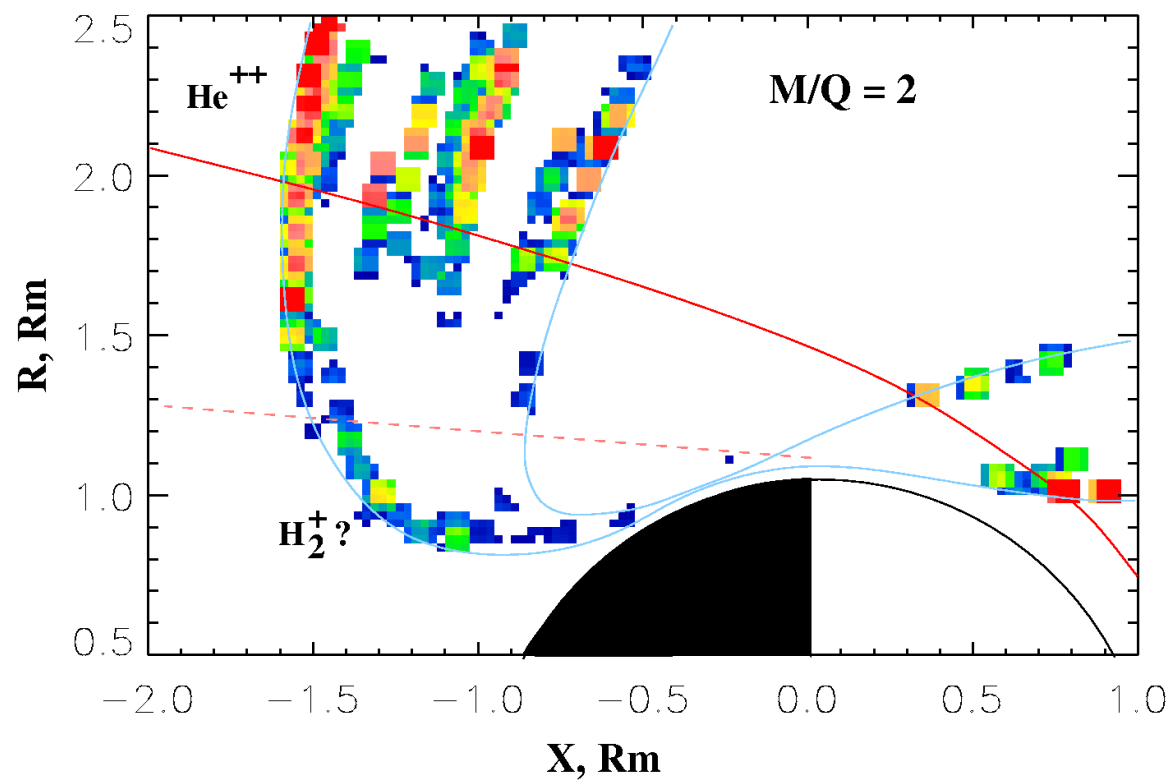


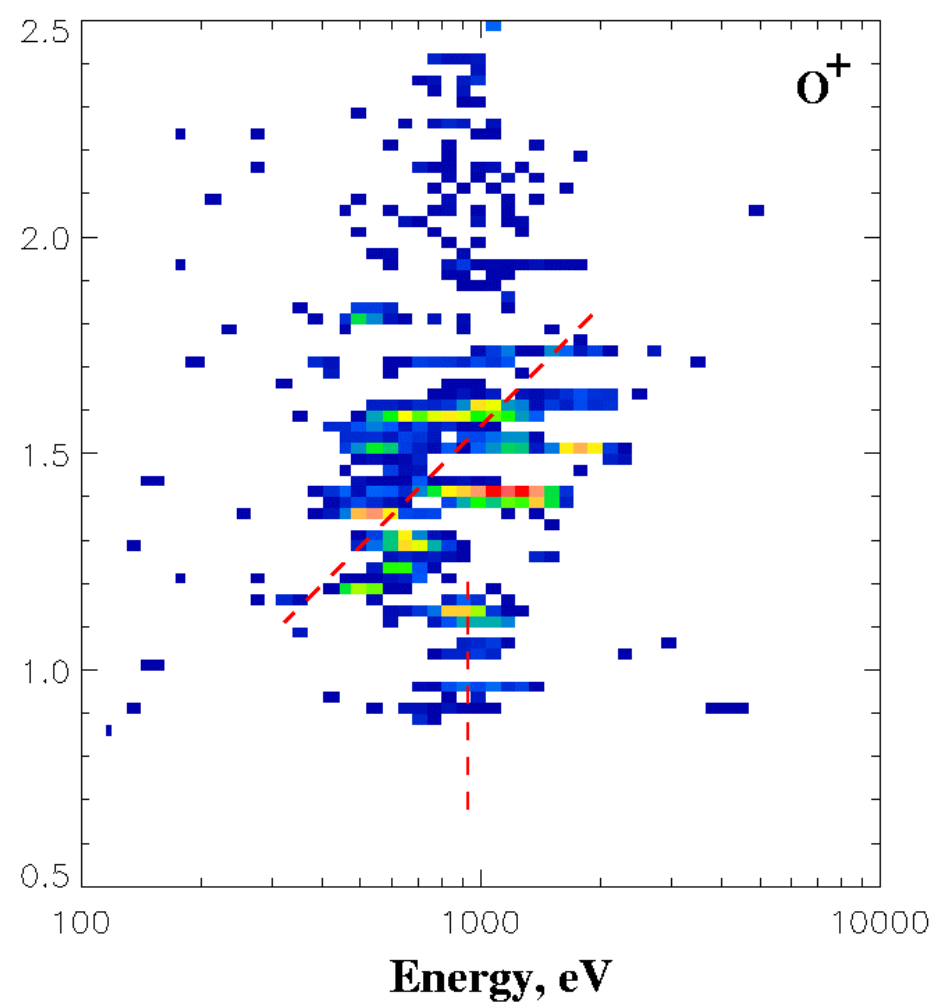
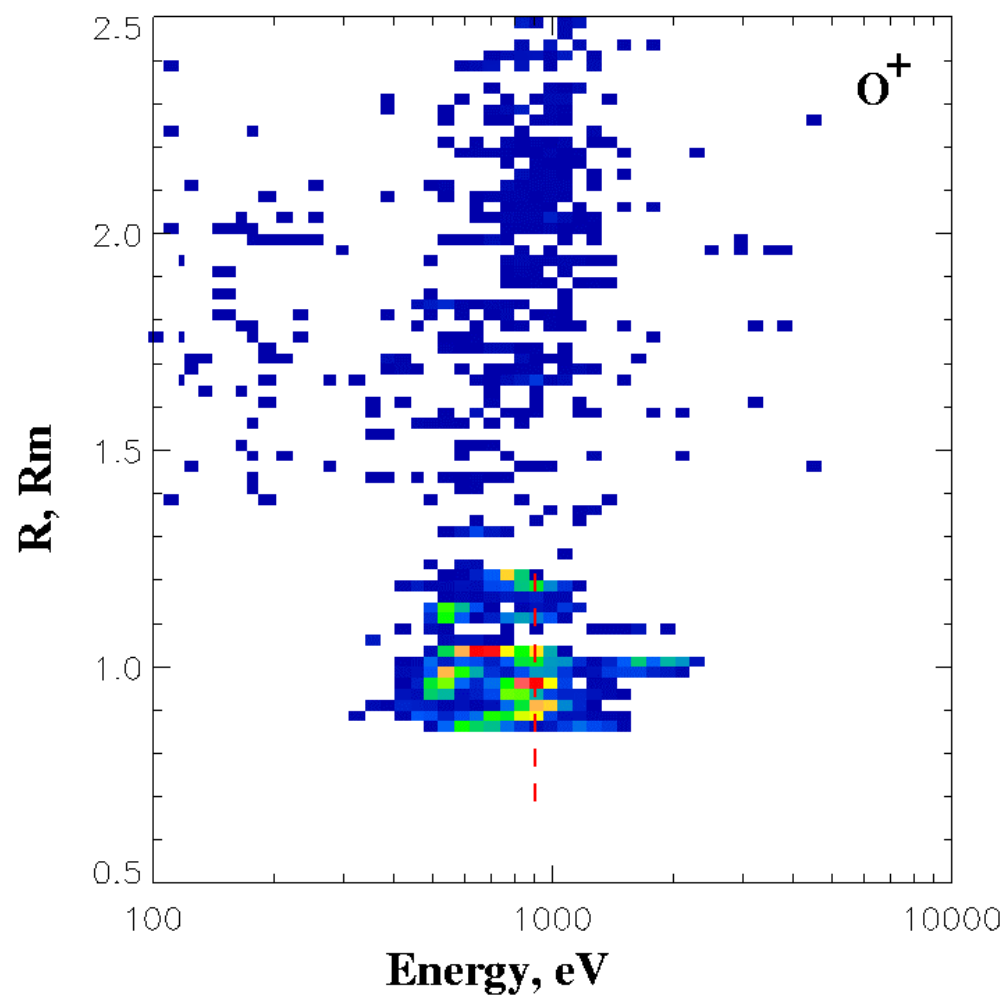


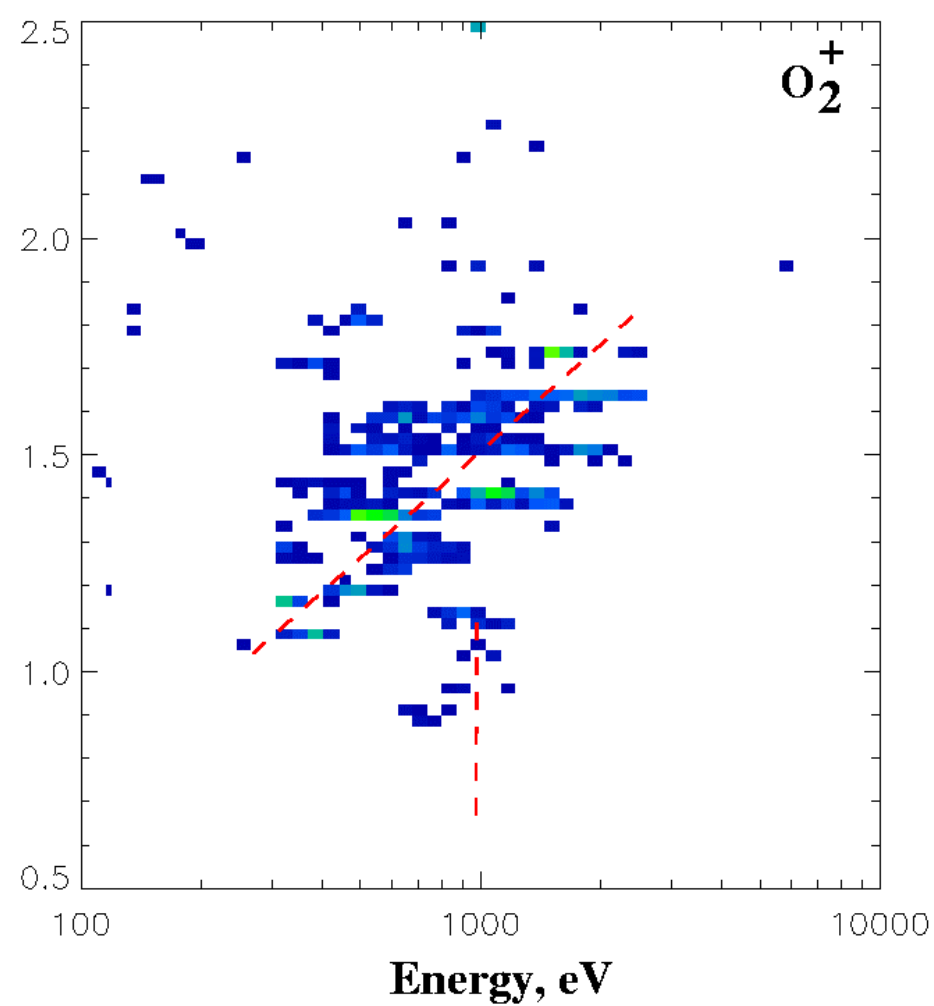
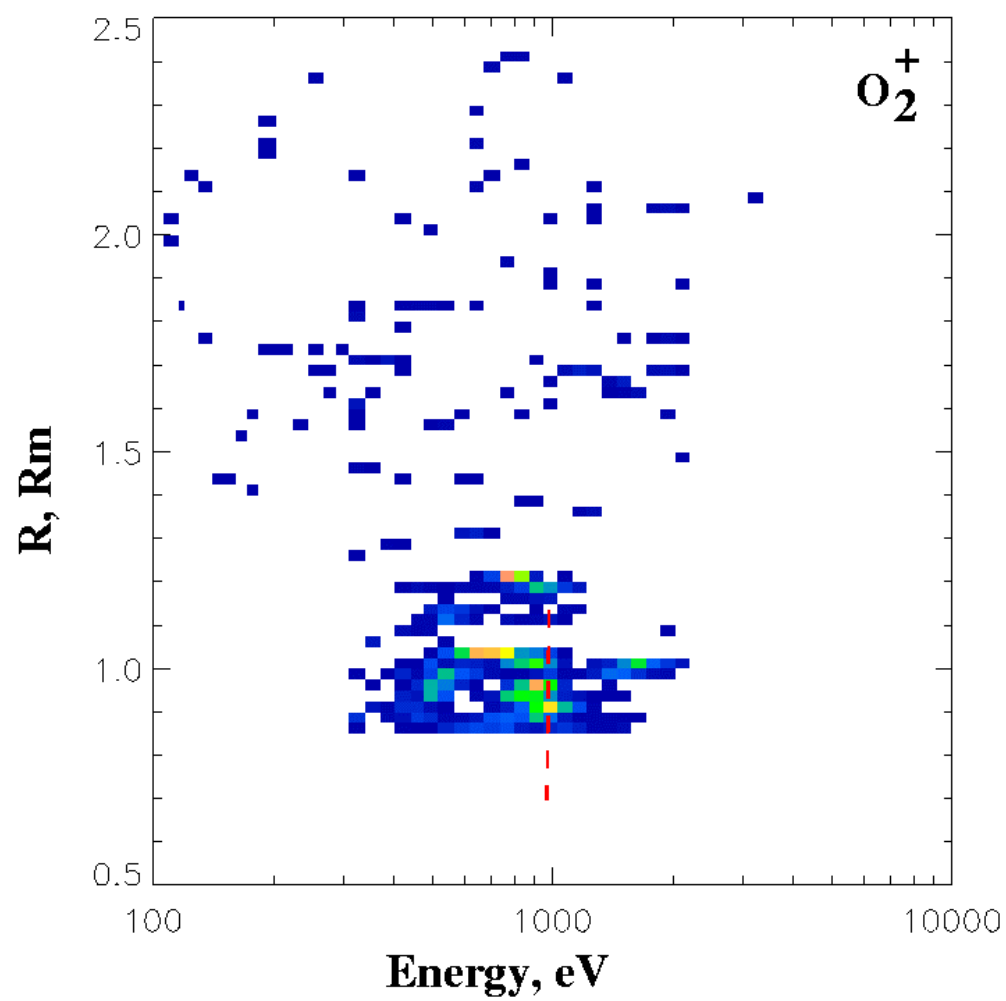












Conlusions:

- There are two regimes of escaping planetary ions :
 - Planet shadow
 - Boundary Layer

Perhaps it depends on IMF_{YZ} direction

- Ions $M/Q = 2$ are seen in the planet shadow

To Do:

IMF reconstruction by MGS observations